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# $P_n$ and $S_n$ tomography across Eurasia to improve regional seismic event locations

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## Abstract

This paper has three motivations: first, to map  $P_n$  and  $S_n$  velocities beneath most of Eurasia to reveal information on a length scale relevant to regional tectonics, second, to test recently constructed 3-D mantle models and, third, to develop and test a method to produce  $P_n$  and  $S_n$  travel time correction surfaces which are the 3-D analogue of travel time curves for a 1-D model. Our third motive is inspired by the need to improve regional location capabilities in monitoring nuclear treaties such as the nuclear Comprehensive Test Ban Treaty (CTBT). To a groomed version of the ISC/NEIC data, we apply the tomographic method of Barmin et al. [Pure Appl. Geophys. (in press)], augmented to include station and event corrections and an epicentral distance correction. The  $P_n$  and  $S_n$  maps are estimated on a  $1^\circ \times 1^\circ$  grid throughout Eurasia. We define the phases  $P_n$  and  $S_n$  as arriving between epicentral distances of  $3^\circ$  and  $15^\circ$ . After selection, the resulting data set consists of about 1,250,000  $P_n$  and 420,000  $S_n$  travel times distributed inhomogeneously across Eurasia. The rms misfit to the entire Eurasian data set from the  $P_n$  and  $S_n$  model increases nearly linearly with distance and averages about 1.6 s for  $P_n$  and 3.2 s for  $S_n$ , but is better for events that occurred on several nuclear test sites and for selected high quality data subsets. The  $P_n$  and  $S_n$  maps compare favorably with recent 3-D models of  $P$  and  $S$  in the uppermost mantle and with recently compiled teleseismic station corrections across the region. The most intriguing features on the maps are the low velocity anomalies that characterize most tectonically deformed regions such as the anomaly across central and southern Asia and the Middle East that extends along a tortuous path from Turkey in the west to Lake Baikal in the east. These anomalies are related to the closing of the Neo-Tethys Ocean and the collision of India with Asia. The uppermost mantle beneath the Pacific Rim back-arc is also very slow, presumably due to upwelling that results from back-arc spreading, as is the Red Sea rift, the Tyrrhenian Sea and other regions undergoing active extension.

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## 1. Introduction

Determination of accurate seismic locations and uncertainties is of prime importance in monitoring the

Comprehensive Nuclear Test Ban Treaty (CTBT). Small magnitude events will only be recorded at a sparse subset of the International Monitoring System (IMS) at regional distances less than  $20\text{--}30^\circ$ . Sparse network locations are subject to significant bias due to regional variations in the structure of the crust and upper mantle. To meet the goals of the CTBT for these small events, this bias must be substantially

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43 reduced in regions of significant structural variability  
44 such as that across much of Eurasia. To do so will  
45 require either a model of the 3-D structures them-  
46 selves or the effects of the structures on the relevant  
47 travel times.

48 This paper has three motivations. The first is to  
49 map  $P_n$  and  $S_n$  velocities beneath most of Eurasia  
50 using regional phase data (viz.,  $P_n$ ,  $S_n$ ) to reveal  
51 information on a scale relevant to regional tectonics.  
52 The second is to test global (e.g., Ekström and  
53 Dziewonski, 1998; Bijwaard et al., 2000; Shapiro et  
54 al., 2000) and regional (e.g., Villaseñor et al., 2001)  
55 3-D seismic models. The third and principal motiva-  
56 tion is to develop and test a method to produce  $P_n$   
57 and  $S_n$  Eurasian travel time correction surfaces. These  
58 surfaces form a common basis for locating seismic  
59 events with regional data alone. Each travel time  
60 correction surface is a map centered on a specific  
61 seismic station. The value at each point on the map is  
62 the travel time observed at the station from a seismic  
63 event located at a specified depth. Usually, the  
64 predicted travel times are presented relative to the  
65 prediction from a 1-D seismic model.

66 The method to estimate  $P_n$  and  $S_n$  that we describe  
67 here is based heavily on earlier efforts by other  
68 researchers (e.g., Hearn et al., 1991). Our method  
69 and earlier incarnations suffer from a number of  
70 problems. These include the fact that  $P_n$  and  $S_n$  are  
71 not monolithic phases that turn at a uniform depth  
72 independent of epicentral distance and tectonic  
73 regime, and it is difficult to separate crustal from  
74 mantle contributions in the observed travel times.  
75 These problems are manifested more strongly on the  
76  $P_n$  and  $S_n$  maps than on the predicted travel time  
77 correction surfaces. Thus, although the methods we  
78 employ may not provide ideal means to estimate  
79 mantle structures, they suffer far fewer problems in  
80 predicting the travel time corrections needed to  
81 improve capabilities to locate regional events. The  
82 results presented here should, therefore, be seen as a  
83 preliminary step toward developing a unified model of  
84 the crust and uppermost mantle that results from  
85 simultaneous inversion of surface wave dispersion  
86 and regional body wave travel times.

87 In the following, we (1) discuss the data set and  
88 the tomographic method, (2) show continental scale  
89 images of  $P_n$  and  $S_n$  variations across Eurasia, (3)  
90 display the resulting travel time correction surfaces

for several IMS or surrogate stations, (4) discuss the  
fit to the regional phase data and inferred uncertain-  
ties in the correction surfaces, and (5) briefly discuss  
some of the velocity anomalies that appear in the  $P_n$   
and  $S_n$  images.

## 2. Data

$P_n$  and  $S_n$  travel times are taken from a groomed  
version of the ISC and NEIC data bases described, in  
part, by Engdahl et al. (1998). ISC travel times are for  
events that occurred from 1964 through 1997 and  
NEIC data are from 1998 to 1999. The locations of  
explosions are replaced with “ground truth” locations  
wherever possible (e.g., Sultanov et al., 1999). We  
define the phases  $P_n$  and  $S_n$  as arriving between  
epicentral distances of  $3^\circ$  and  $15^\circ$ .  $P_n$  and  $S_n$  may  
dip into the mantle substantially, particularly beyond  
epicentral distances of  $\sim 8^\circ$ . The depth of penetration  
will depend on the vertical derivative of velocity,  
which will vary spatially. The truncation of the data  
set to include rays only if epicentral distances are less  
than  $10\text{--}12^\circ$ , as in some other studies (e.g., Hearn  
and James, 1994), would severely restrict path cover-  
age in some areas of Eurasia. To utilize longer paths, it  
is desirable to correct for the effect of ray penetration  
into the uppermost mantle. We discuss this correction  
in Section 3.

This data set consists of 3,672,268  $P_n$  phases and  
1,346,676  $S_n$  phases for 5418 stations and 149,929  
events worldwide. Data are used in the inversion if the  
residual relative to the prediction from the 1-D model  
ak135 (Kennett et al., 1995) is less than 7.5 s for  $P$   
and 15 s for  $S$ , if the event depth is within the crust or  
less than 50 km deep, if the azimuthal gap to all  
reporting stations for the event is less than  $180^\circ$ , and  
if the nominal error ellipse is less than  $1000\text{ km}^2$  in  
area. These selection criteria reduce the data set to  
1,636,430  $P_n$  and 493,734  $S_n$  phases worldwide. Most  
of these paths cross Eurasia ( $0\text{--}80^\circ\text{N}$  latitude and  
 $-10^\circ\text{W}\text{--}180^\circ\text{E}$  longitude plus a buffer zone), the  
numbers being 1,257,052 for  $P_n$  and 422,634 for  $S_n$ .  
Data density is shown in Fig. 1. Because path lengths  
for these phases are by definition short ( $< 15^\circ$ ), paths  
only exist in regions where both sources and receivers  
are common. Thus, the path distribution is highly  
heterogeneous across the region.

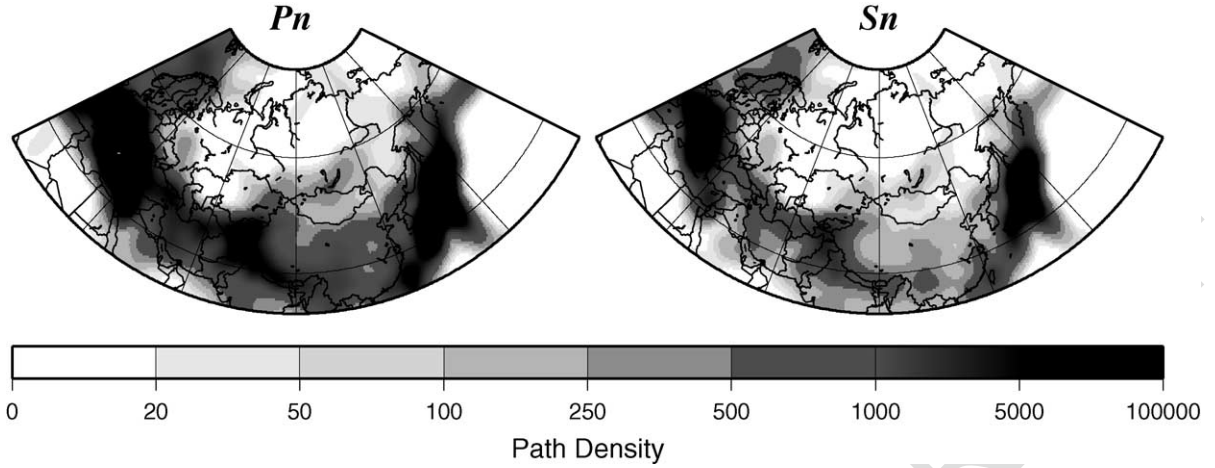


Fig. 1. Path density for the  $P_n$  and  $S_n$  data, defined as the number of paths intersecting a  $2^\circ \times 2^\circ$  cell ( $\sim 50,000 \text{ km}^2$ ).

136 We further reduce this data set by rejecting late  
 137 arriving travel times at epicentral distances from  $3^\circ$   
 138 to  $6^\circ$  that may be misidentified crustal phases (e.g.,  
 139  $P_g$ ) or Moho reflections ( $PmP$ ,  $SmS$ ). In addition, in  
 140 the tomography, we reject measurements misfit by  
 141 the starting model at more than  $2\sigma$ , where  $\sigma$  is the  
 142 average misfit produced by the starting model. This is  
 143 done to help stabilize the station and event correc-  
 144 tions. However, we report misfit statistics relative to  
 145 the entire data set across Eurasia.

### 146 3. Method

147 The observed travel time,  $t_{\text{obs}}$ , is modeled as  
 148 follows:

$$t_{\text{obs}} = t_m + t_{\text{crust\_sta}} + t_{\text{crust\_evt}} + \delta t_{\text{sta}} + \delta t_{\text{evt}} + \delta t(\Delta) + \delta t_m, \quad (1)$$

149 where  $t_m$  is the predicted travel time for rays  
 151 through the mantle part of the input reference  
 152 model, the contributions to the travel time due to  
 153 the crustal part of the reference model on the event  
 154 and station sides are  $t_{\text{crust\_sta}}$  and  $t_{\text{crust\_evt}}$ , the  
 155 station and event delays or static corrections are  
 156  $\delta t_{\text{sta}}$  and  $\delta t_{\text{evt}}$ ,  $\delta t(\Delta)$  is the distance correction,  $\delta t_m$   
 157 is the travel time correction for the mantle part of  
 158 the path, and  $\delta$  is epicentral distance. Thus,  $t_m$ ,

$t_{\text{crust\_sta}}$ , and  $t_{\text{crust\_evt}}$  are predicted by the reference  
 model and  $\delta t_{\text{sta}}$ ,  $\delta t_{\text{evt}}$ ,  $\delta t(\Delta)$ , and  $\delta t_m$  are estimated.  
 If  $v_m$  is the velocity along path  $p$  in the reference  
 model and  $\delta v_m$  is the model perturbation along the  
 same path, then

$$t_m = \int_p \frac{ds}{v_m}, \quad (2)$$

$$\delta t_m = - \int_p \frac{\delta v_m}{v_m^2} ds. \quad (3)$$

We assume that the ray through the perturbed  
 model,  $v_m + \delta v_m$ , takes the same path as the ray  
 through the reference model. In practice, we esti-  
 mate the 2-D quantity  $\delta v_m$  from which we compute  
 $\delta t_m$  for each ray  $p$ .

We use CRUST5.1 (Mooney et al., 1998) as the  
 starting (reference) model in the crust and for  
 mantle  $P$  and  $S$ . At each geographical point,  
 CRUST5.1 only has one value of  $P$  and one value  
 of  $S$  for the mantle, intended to characterize the  
 velocity immediately below Moho. Thus, for the  
 reference model, the mantle leg of each path  $p$  is  
 horizontal, following directly below Moho as shown  
 in Fig. 2. This is, in fact, a common approximation  
 in  $P_n$  and  $S_n$  tomography, but realistic rays dive into  
 the mantle to depths that depend on a nonzero  
 vertical velocity derivative as Fig. 3b shows. Fig.

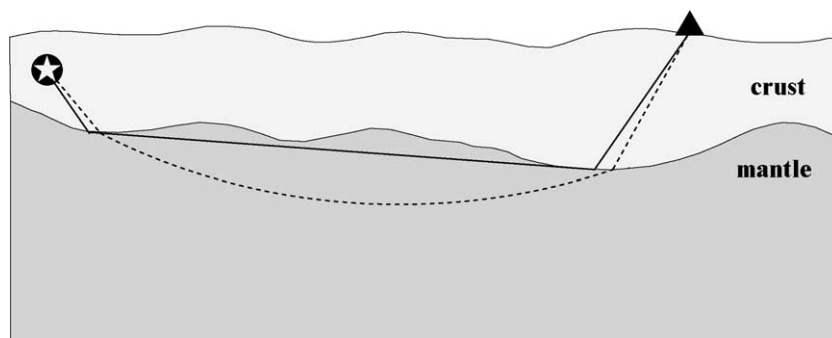


Fig. 2. Illustration of sources of error in the tomography. Real ray paths (dashed line) follow different paths through both the crust and mantle than the hypothesized rays (solid line) used in tomography. In particular, real paths dip deeper into the mantle as epicentral distances increase in a way that depends on the vertical velocity gradient.

184 3c attempts to quantify the error made by the  
 185 horizontal ray approximation, by comparing the  
 186 travel times diving into the mantle through a recent  
 187 Eurasian 3-D model (Shapiro et al., 2000) with  
 188 those computed for a model in which the rays  
 189 propagate horizontally directly beneath Moho. The  
 190 horizontal ray approximation produces an error that  
 191 is a relatively smooth function of distance. For most  
 192 of the continent, the estimated errors are similar and  
 193 grade smoothly to a travel time error predicted to be  
 194 about  $-2.5$  s at  $15^\circ$  for  $P$  velocities. This moti-  
 195 vates the introduction into Eq. (1) of a term that is a  
 196 smooth function of distance, which we call the  
 197 distance correction,  $\delta t(\Delta)$ . The correction  $\delta t(\Delta)$   
 198 attempts to reduce the mantle velocities distributed  
 199 in 3-D to a single 2-D datum surface which, by  
 200 design, lies directly below Moho. The paths from  
 201 WMQ and ANTO in Western China and Turkey,  
 202 respectively, exhibit anomalously high and low  
 203 vertical velocity gradients. In these regions, the  
 204 errors produced by the horizontal ray approximation  
 205 will be atypical and will be poorly corrected by  
 206  $\delta t(\Delta)$ .

207 Although the distance correction is an average  
 208 across the continent, it allows us to fit data over a  
 209 broader distance range than would be possible without  
 210 the correction. We find that with this correction, the  
 211 tomographic maps agree well with those produced  
 212 with short path data alone (epicentral distances less  
 213 than  $10^\circ$ ) in those regions where tomographic maps  
 214 can be constructed reliably using only the short path  
 215 data. The use of a 3-D model to compute the distance

correction is beyond the scope of the present work,  
 but is an important direction for future research.

We follow Hearn and collaborators (e.g., Hearn  
 and Clayton, 1986; Hearn et al., 1991; Hearn and  
 James, 1994; and elsewhere) and estimate event and  
 station corrections. These corrections are designed  
 to compensate for errors in the reference crustal  
 model, errors in the prediction of the location of the  
 mantle piercing points, and errors in event locations  
 and origin times. A correction is estimated for a  
 station if there are phase picks from at least seven  
 events made at that station and an event correction  
 is estimated for all events for which there are at  
 least 20 reporting stations. The asymmetry in this  
 condition is because there are more physical phe-  
 nomena modeled with the event correction than  
 with the station correction (e.g., mislocation, origin  
 time error). The station and event corrections are  
 undamped.

As presented here, the  $P_n$  and  $S_n$  maps are defined  
 over a two-dimensional surface and, therefore, may be  
 estimated with the same 2-D tomographic method we  
 developed for surface wave tomography (Barmin et  
 al., in press). The inversion for  $P_n$  and  $S_n$  is approx-  
 imately the same except that from the reference  
 model, we compute source and receiver side Moho  
 penetration points and use these points as the starting  
 and ending points of the ray during inversion (see Fig.  
 2). The approximation comes from the assumption  
 that  $P_n$  and  $S_n$  rays are “horizontal” in a spherical  
 mantle and propagate directly below Moho, as dis-  
 cussed above and depicted in Fig. 2.

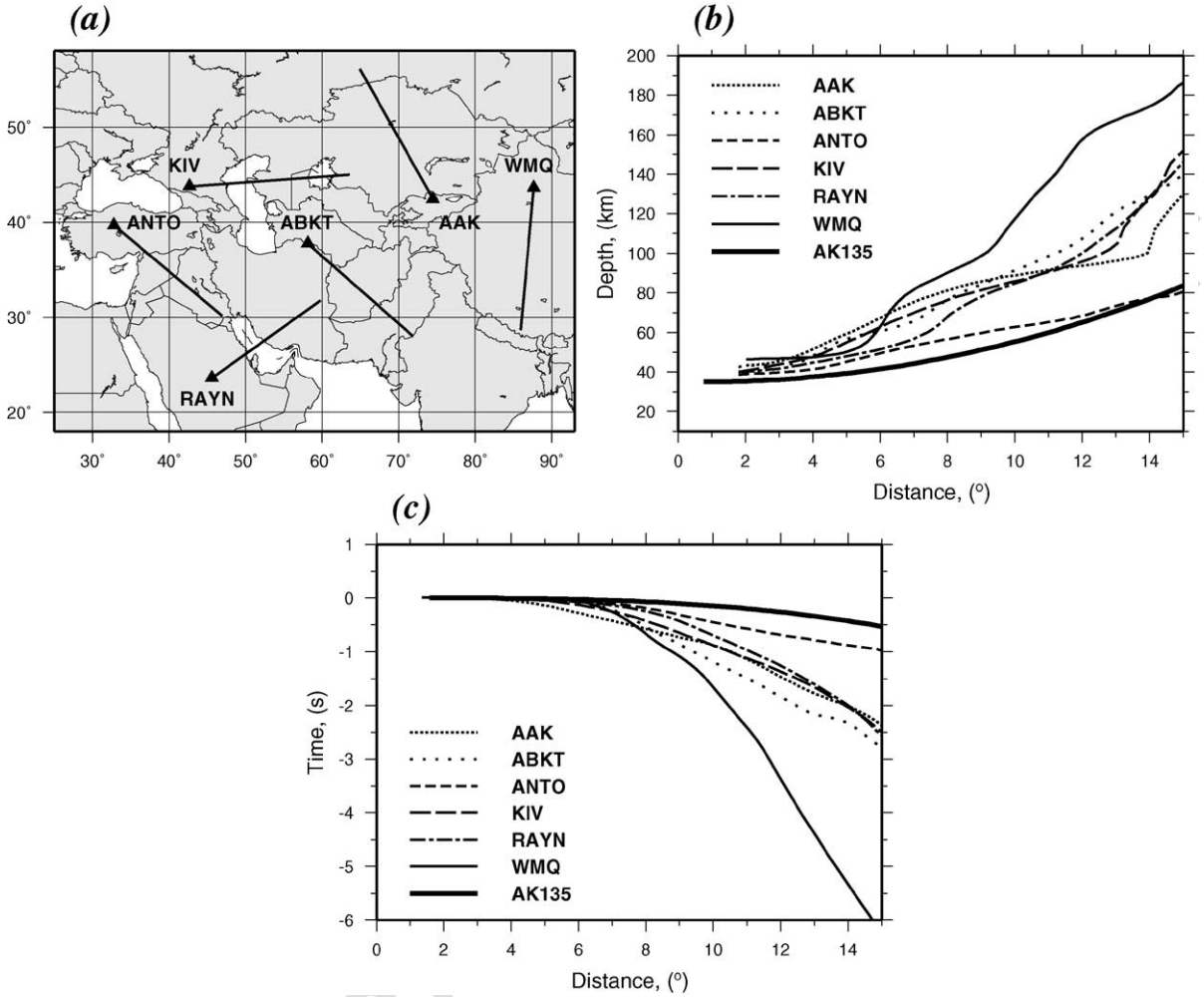


Fig. 3. Examples to justify the distance correction,  $\delta t(\Delta)$ . (a) The location of the six 2-D velocity profiles used in (b) and (c). Each profile starts from a seismic station and runs for 15°. The 3-D model used is that of Shapiro et al. (2000). (b) Turning point curves for the six profiles in (a) and for the 1-D model ak135. (c) Each curve is the difference between the travel time travel computed through the 3-D model of Shapiro et al. (2000) using the ray shooting method of Cerveny and Psencik (1988) and the travel time through the same model with a horizontal ray. This provides an estimate of the error in P-wave travel time caused by the horizontal ray approximation for the six profiles in (a).

248 In the method of Barmin et al. (in press), the model  
 249 is constructed on an equally spaced grid such that the  
 250 following figure-of-merit is minimized:

$$\begin{aligned}
 & (\mathbf{Gm} - \mathbf{d})^T \mathbf{C}^{-1} (\mathbf{Gm} - \mathbf{d}) + \sum_{k=0}^n \alpha_k^2 \|F_k(\mathbf{m})\|^2 \\
 & + \sum_{k=0}^n \beta_k^2 \|H_k(\mathbf{m})\|^2, \quad (4)
 \end{aligned}$$

which is a linear combination of data misfit, model 252  
 roughness, and the amplitude of the perturbation to a 253  
 reference model. The vector  $\mathbf{m}$  represents the esti- 254  
 mated model,  $\delta v_m$ , which is a perturbation relative to a 255  
 reference across the region of interest,  $\mathbf{G}$  is the 256  
 forward operator that computes travel time from the 257  
 estimated model,  $\mathbf{d}$  is the data vector,  $\mathbf{C}$  is the data 258  
 covariance matrix or matrix of data weights,  $F$  is a 259  
 Gaussian smoothing operator, and  $H$  is an operator 260  
 that penalizes the norm of the model  $\mathbf{m}$  in regions of 261

262 poor data coverage. The spatial smoothing operator is  
 263 defined over the 2-D model as follows

$$F_k(\mathbf{m}) = m_k(\mathbf{r}) - \int_S S_k(\mathbf{r}, \mathbf{r}') m_k(\mathbf{r}') d\mathbf{r}', \quad (5)$$

264 where  $S_k$  is a smoothing kernel:

$$S_k(\mathbf{r}, \mathbf{r}') = K_{0k} \exp\left(-\frac{|\mathbf{r} - \mathbf{r}'|^2}{2\sigma_k^2}\right) \quad (6)$$

$$\int_S S_k(\mathbf{r}, \mathbf{r}') d\mathbf{r}' = 1, \quad (7)$$

269 and  $\sigma_k$  is the spatial smoothing width or correlation  
 270 length. The minimization of the expression in Eq. (5)  
 271 explicitly ensures that the estimated model approx-

272 imates a smoothed version of the model. The maps are  
 273 estimated on a  $1^\circ \times 1^\circ$  grid across Eurasia with  
 274  $\sigma_k = 150$  km.

275 We refer to the  $P_n$  and  $S_n$  maps together with  
 276 the parametric corrections  $\delta t_{sta}$ ,  $\delta t_{evt}$ , and  $\delta t(\delta)$  as  
 277 the CU  $P_n/S_n$  model to distinguish it from our  
 278 recent 3-D models (e.g., Villaseñor et al., 2001;  
 279 Shapiro et al., 2000). To compute travel time  
 280 correction surfaces, the distance correction, the  
 281 station delays, and the  $P_n$  or  $S_n$  map must be used.

282 The method of Barmin et al. (in press) allows us  
 283 to estimate  $2\psi$  and  $4\psi$  azimuthal anisotropy simulta-  
 284 neously with isotropic  $P_n$  and  $S_n$ . We find, however,  
 285 that the estimates of azimuthal anisotropy are not  
 286 robust with respect to data subsetting and variations

## Station & Event Corrections

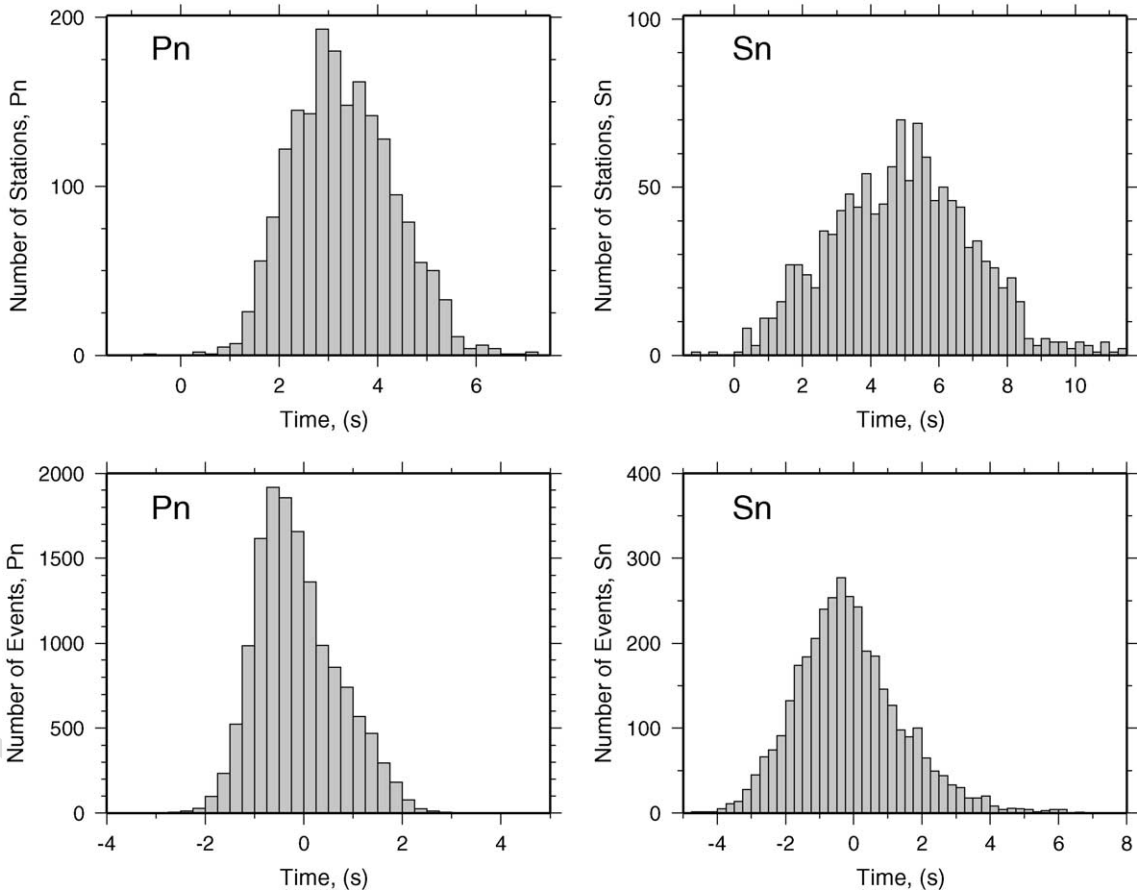


Fig. 4.  $P_n$  and  $S_n$  station (top row) and event (bottom row) corrections: histograms of values.



287 in damping. In addition, the joint inversion for  
 288 isotropic and anisotropic structures dominantly  
 289 affects only the amplitudes of the isotropic maps,  
 290 but no more so than the choice of isotropic damping  
 291 which is itself largely arbitrary. Consequently, we  
 292 report only isotropic  $P_n$  and  $S_n$  maps here and safely  
 293 can ignore the effects of azimuthal anisotropy on  
 294 these estimates.

#### 295 4. $P_n$ and $S_n$ tomography

296 We estimate station delays  $\delta t_{\text{sta}}$ , event delays  
 297  $\delta t_{\text{evt}}$ , the distance correction curve  $\delta t(\Delta)$ , and the

298 2-D tomographic quantity  $\delta v_m(\theta, \varphi)$  which repre-  
 299 sents lateral variations in seismic velocities in the  
 300 uppermost mantle. Latitude and longitude are  $\theta$  and  
 301  $\varphi$ , respectively. There are strong and, essentially,  
 302 unresolvable trade-offs between subsets of these  
 303 quantities. For example, a constant velocity shift  
 304 in the uppermost mantle could be fit either by a  
 305 constant shift in  $\delta v_m$  or by introducing a linear  
 306 trend in  $\delta t(\Delta)$ . The station delays,  $\delta t_{\text{sta}}$ , also  
 307 strongly trade-off with  $\delta v_m$  directly beneath the  
 308 station and the value of the distance correction at  
 309 an epicentral distance of  $3^\circ$ . The estimated values  
 310 depend strongly on the inversion algorithm and the  
 311 order in which the corrections are estimated, if the

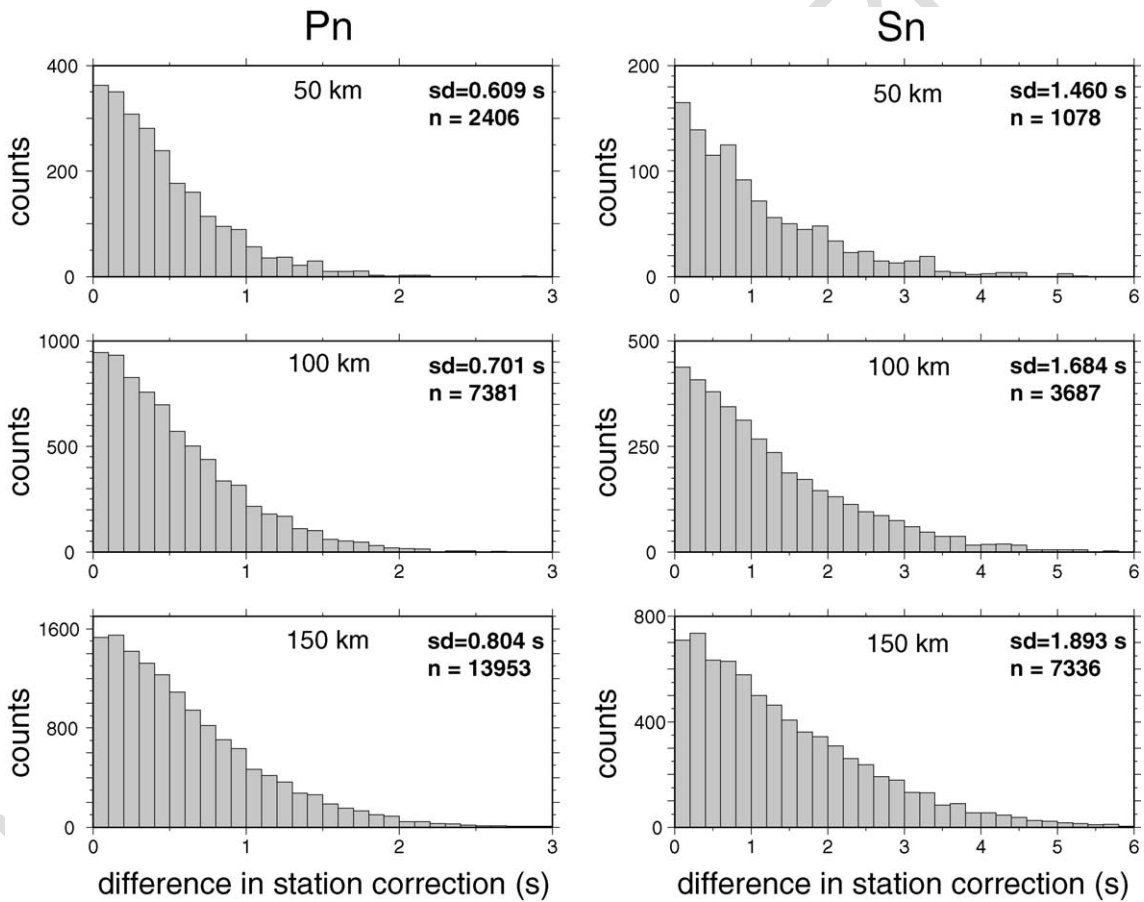


Fig. 5. Spatial coherence of station corrections. Each plot shows the distribution of the absolute value of the difference between the station corrections for nearby stations. In the upper row, the pair of stations lies within 50 km of one another, 100 km in the middle row, and 150 km in the bottom row. Results for  $P_n$  are in the left column and for  $S_n$  in the right column. The standard deviation (S.D.) of each distribution is marked on each plot as is the number of station pairs ( $n$ ).

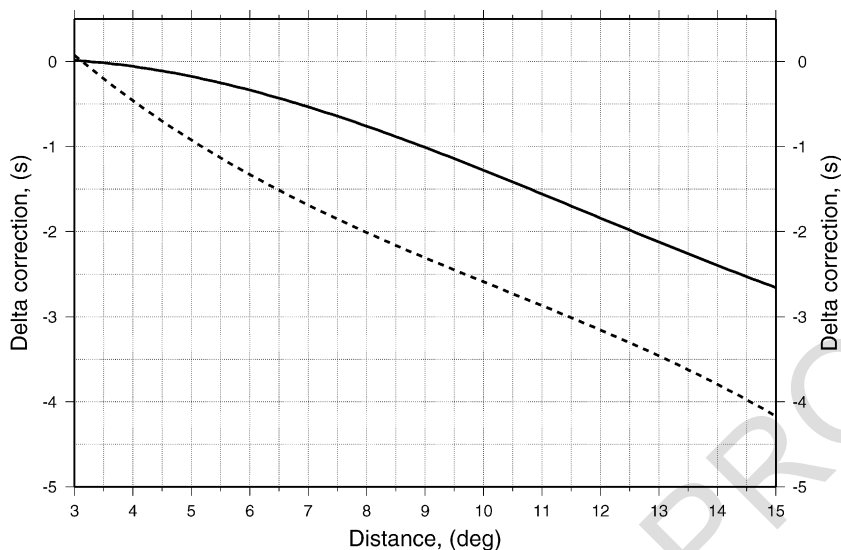


Fig. 6. Estimated distance correction,  $\delta t(\Delta)$ , for  $P_n$  (solid line) and  $S_n$  (dashed line). The distance corrections are constrained to be approximately zero at an epicentral distance of  $3^\circ$ .

312 process is iterative rather than simultaneous. If the  
 313 process were simultaneous, then values would  
 314 depend on the relative weights assigned to each  
 315 correction. We constrain  $\delta t(\Delta)$  to be approximately  
 316 zero at  $3^\circ$  and let the curve  $\delta t(\Delta)$  have only a  
 317 moderate negative slope. Thus, we choose to fit  
 318 much of the signal with station delays and allow a  
 319 substantial constant shift in  $\delta v_m$ .

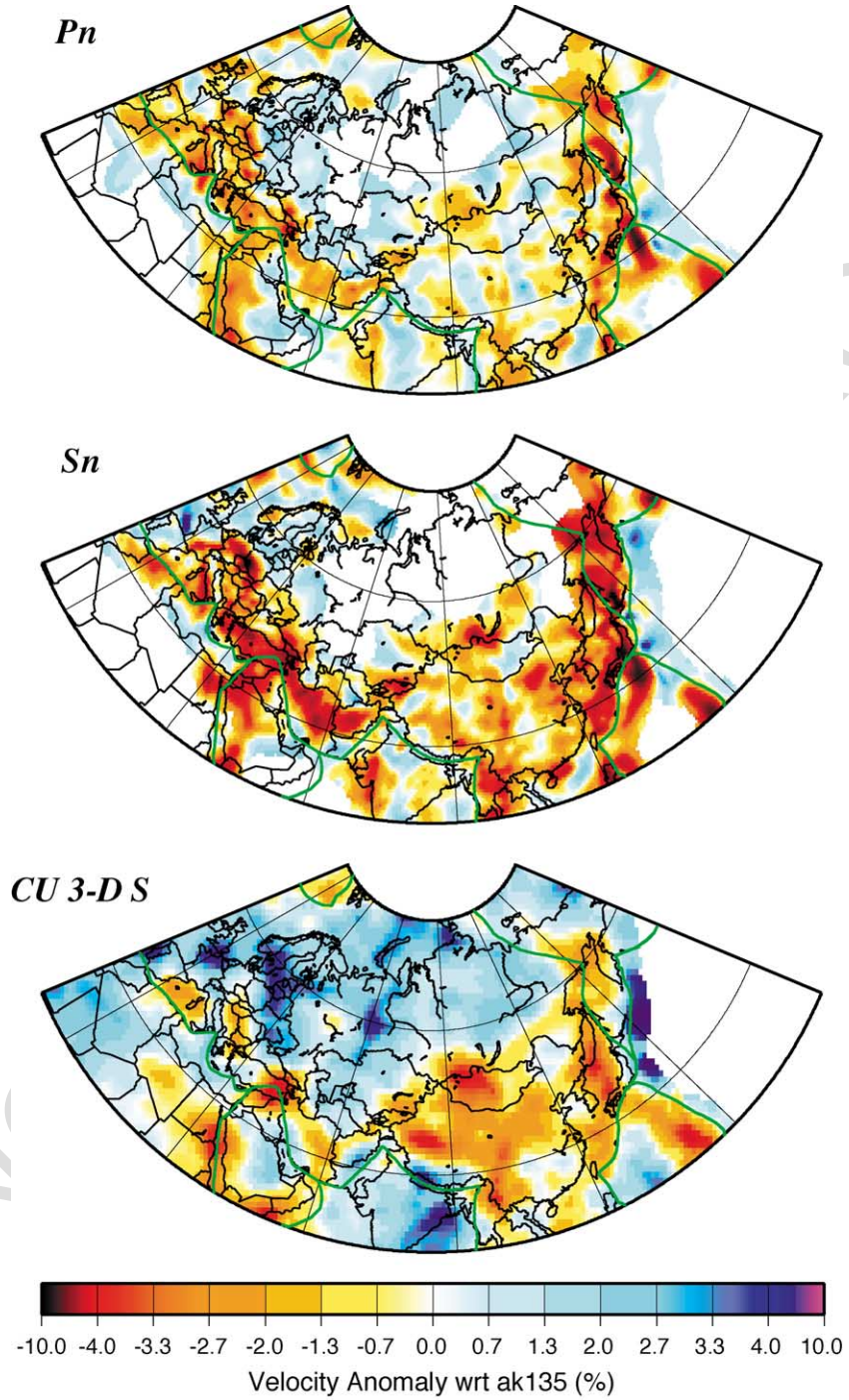
320 The magnitudes of station and event delays,  $\delta_{sta}$  and  
 321  $\delta_{evt}$ , are summarized in Fig. 4. Not surprisingly, the  $S_n$   
 322 corrections are typically larger than those for  $P_n$ ,  
 323 presumably because  $S$  variations in the crust and upper  
 324 mantle are typically larger than  $P$  by about a factor of  
 325 two. The  $P_n$  and  $S_n$  station delays are geographically  
 326 coherent and correlate with one another with a poorly  
 327 determined  $S/P$  ratio of about 1.8 relative to the mean of  
 328 each distribution. Fig. 5 shows that if two stations are  
 329 closer than 50 km apart, the standard deviation of their  
 330 station delay is about 0.6 s for  $P_n$  and 1.5 s for  $S_n$ . Part of  
 331 this difference is structural, as differences in the station  
 332 corrections grow with the separation between the  
 333 stations. Stations and events are not uniformly distrib-

uted over the continent, with stations predominantly in  
 stable continental regions and events in tectonically  
 deformed regions. For this reason, together with the  
 fact that the delays are taken relative to a model, the  
 delays are not expected to be zero-mean and, in fact,  
 display a positive mean for the stations and a negative  
 mean for the events.

Although we estimate  $P_n$  corrections only for  
 about 60% of the stations and half of the events  
 worldwide and for  $S_n$  the numbers are about 50%  
 and 25%, respectively, the great majority of the  
 measurements emanate from events and are re-  
 corded at stations that have corrections. This is  
 particularly true for  $P_n$ , where only about 3% of  
 the measurements are made at stations without  
 corrections and 12% are for events without event  
 corrections. For  $S_n$ , the numbers are 4% and 33%,  
 respectively. Thus, most measurements have the full  
 complement of corrections applied. For stations and  
 events for which we have not estimated corrections,  
 we set the corrections equal to the mean of the  
 distributions shown in Fig. 4.

Fig. 7.  $P_n$  and  $S_n$  maps estimated across Eurasia. Values are relative to the prediction from ak135 at the top of the mantle, 8.04 km/s for  $P_n$  and 4.48 km/s for  $S_n$ . The bottom map is the equivalent isotropic  $S$  velocity at the top of the mantle from a recently estimated model beneath Eurasia (e.g., Villaseñor et al., 2001). Regions in which path density is less than about 20 paths/50,000 km<sup>2</sup> are poorly constrained by the data, and these regions are shaded white.





356 The distance correction is shown in Fig. 6. The  
 357 shape of the  $P$  distance correction is different from  
 358 that predicted by Fig. 3, but the value of the correc-

tion at  $15^\circ$  is about the same ( $\sim -2.5$  s). There is, in  
 addition, a constant offset in  $\delta v_m$  equal to about  
 $-100$  m/s relative to the average  $S_n$  velocity of

359  
 360  
 361

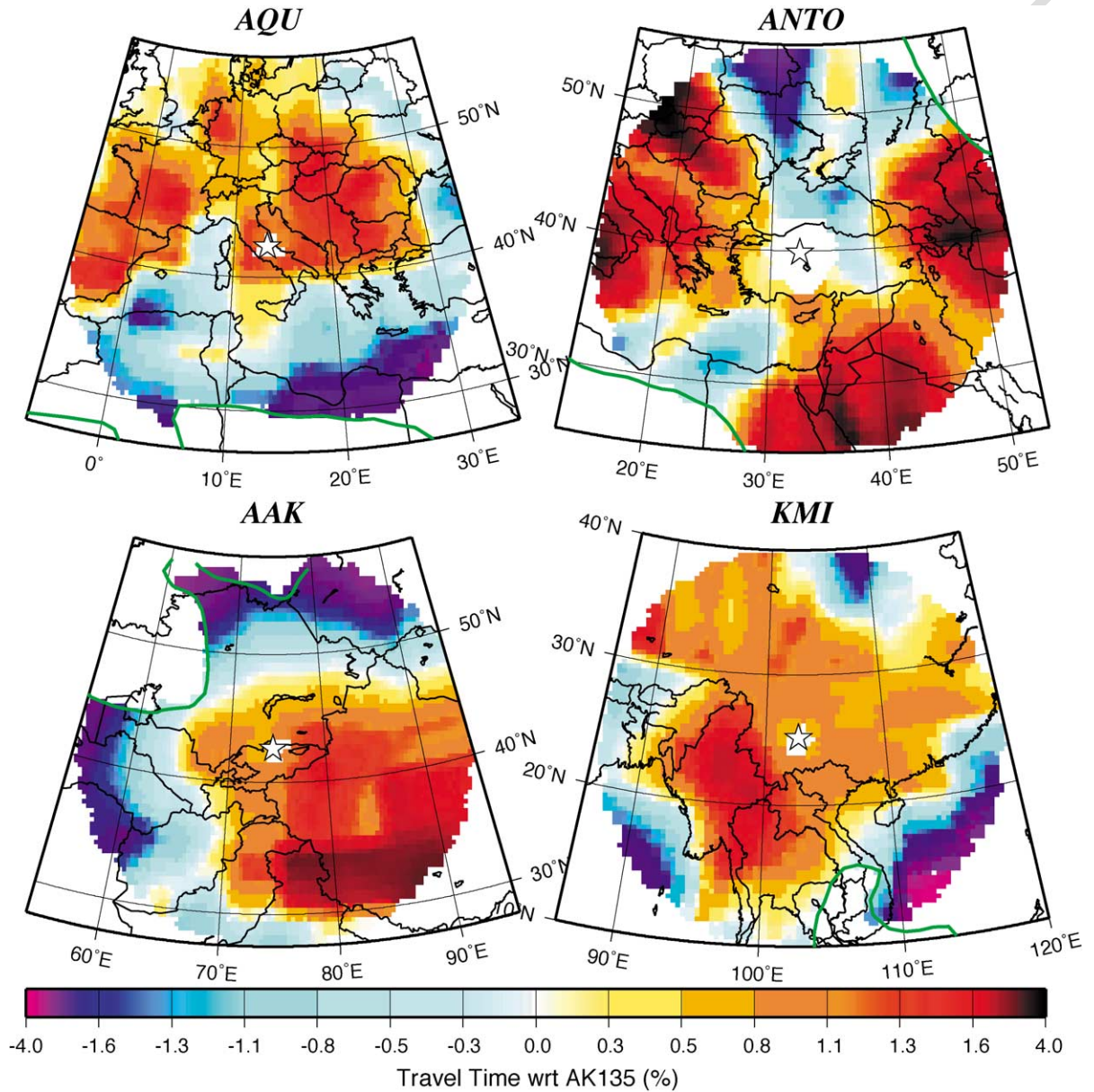


Fig. 8. Travel time correction surfaces for  $P_n$  for four IMS stations or surrogates identified with stars (AQU, L'Aquila Italy; ANTO, Ankara Turkey; AAK, Ala-Archa Kyrghyzstan; KMI, Kunming China). These surfaces exist only where  $P_n$  data density is locally greater than 20 paths/50,000 km<sup>2</sup>, marked by the green contour, and to  $15^\circ$  from the station. Low data density regions are shaded white. The blocky features that appear in the AAK surface are remnants of the  $5^\circ \times 5^\circ$  starting model CRUST5.1.

362 CRUST5.1. We have greater confidence in the deci-  
 363 sions we reached to resolve the trade-off between  $\delta v_m$   
 364 and  $\delta t(\delta)$  for  $P$  than for  $S$ .

365 The estimated  $P_n$  and  $S_n$  maps are shown in Fig. 7.  
 366 Because our tomographic method penalizes the ampli-  
 367 tude of the maps in regions of poor data coverage and  
 368 the estimated maps are perturbations to a reference  
 369 state, the maps revert to the reference model where data  
 370 coverage is poor, i.e., less than 15–20 paths for each  
 371  $2^\circ \times 2^\circ$  cell. The areas of poor data coverage are  
 372 identified as white regions in Fig. 7. Both the  $P_n$  and

$S_n$  maps demonstrate poor coverage across the shield 373  
 and platform regions of northern Russia and Kazakh- 374  
 stan, in the oceans, and across North Africa. Elsewhere, 375  
 the spatial resolution of the maps is estimated to be 376  
 between 150 and 300 km. 377

The  $P_n$  and  $S_n$  anomalies in Fig. 7 are highly 378  
 correlated and differ in amplitude by about a factor 379  
 of two, such that  $\delta v_s/v_s \sim 2\delta v_p/v_p$ . The anomalies 380  
 also compare well with known tectonic features and 381  
 with the patterns of velocity variations at the top of 382  
 the upper mantle in the  $2^\circ \times 2^\circ$  3-D shear velocity 383

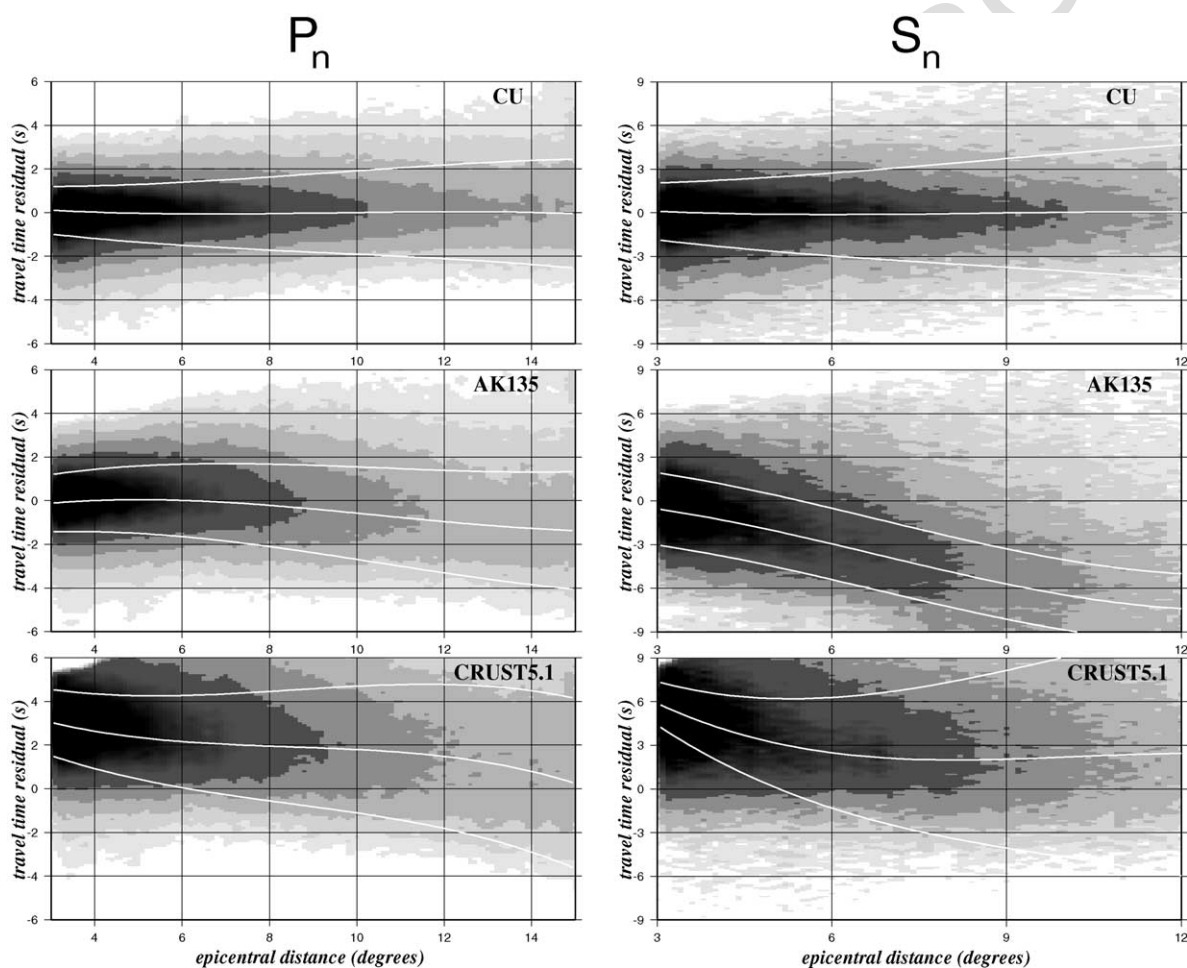


Fig. 9. Shaded plots of the  $P_n$  and  $S_n$  travel time residuals (observed–predicted) for the Eurasian data set presented versus epicentral distance. Results for three models are shown: (top) the CU  $P_n/S_n$  model, (middle) the 1-D model ak135 (Kennett et al., 1995), and (bottom) the laterally heterogeneous crustal,  $P_n$ , and  $S_n$  model CRUST5.1 (Mooney et al., 1998). Darker shades indicate a larger number of residuals and the white lines show the smoothed local mean and  $\pm 1\sigma$ . Overall means and standard deviations are summarized in Table 1 and distance trends are shown in Fig. 10.



t1.1 Table 1

Summary of misfits to the whole Eurasian data set displayed in Fig. 9

Model	$P_n$		$S_n$	
	Mean (s)	$\sigma$ (s)	Mean (s)	$\sigma$ (s)
CU $P_n/S_n$	0.01	1.63	0.03	3.20
ak135	-0.01	1.88	-2.21	4.29
CRUST5.1	2.51	2.14	4.26	4.25

384 model of Villaseñor et al. (2001). Smaller scale  
 385 features, however, are apparent in the  $P_n$  and  $S_n$   
 386 maps presented here and the amplitudes of the  $S_n$   
 387 map are somewhat larger than in the 3-D  $S$  model.  
 388 Villaseñor et al. (2001) also show that the anoma-  
 389 lies are similar to those in the teleseismic  $P$  model  
 390 of Bijwaard et al. (1998) and Engdahl and Ritz-  
 391 woller (2001) demonstrate that the anomalies corre-  
 392 late with teleseismic station corrections. Thus, the  
 393 patterns of high and low velocities are robust and  
 394 are apparent in a number of different data sets at  
 395 both regional and global scales.

396 A thorny problem arises in comparing model  
 397 predictions; namely, the transversely isotropic nature  
 398 of  $S$  models in the upper mantle. Regional  $S$  is a split  
 399 phase and it is unclear if, on average, the observed  
 400 travel times correspond to SH, SV or some linear  
 401 combination. We show in Fig. 7 an “equivalent  
 402 isotropic” shear velocity computed from our 3-D  $S$   
 403 model, which is approximately the average of  $v_{sh}$   
 404 and  $v_{sv}$ . Any offset between the predictions from this  
 405  $S$  model and the estimated  $S_n$  map may result from  
 406 anisotropy or from improperly resolving the trade-off  
 407 between the distance correction and  $\delta v_m$ , as dis-  
 408 cussed above. Therefore, it may be most reasonable  
 409 to compare variations around a poorly determined  
 410 mean, although Villaseñor et al. (2001) show that  
 411 inhomogeneities in  $v_{sh}$  and  $v_{sv}$  in the upper mantle  
 412 are not correlated everywhere.

## 413 5. Travel time correction surfaces

414 Travel time correction surfaces are a computa-  
 415 tional convenience commonly used for locating seis-  
 416 mic events with regional data alone. Each correction  
 417 surface is a map centered on a specific seismic  
 418 station. The value at each point on the map is the

travel time predicted at the station from a seismic  
 event at a specified depth. They are, therefore, the  
 analogue for a 3-D model of travel time curves for 1-  
 D models. Usually, the predicted travel times are  
 presented relative to the prediction from a 1-D  
 seismic model. The accuracy of regional event loca-  
 tions will depend directly on the accuracy of the  
 correction surfaces.

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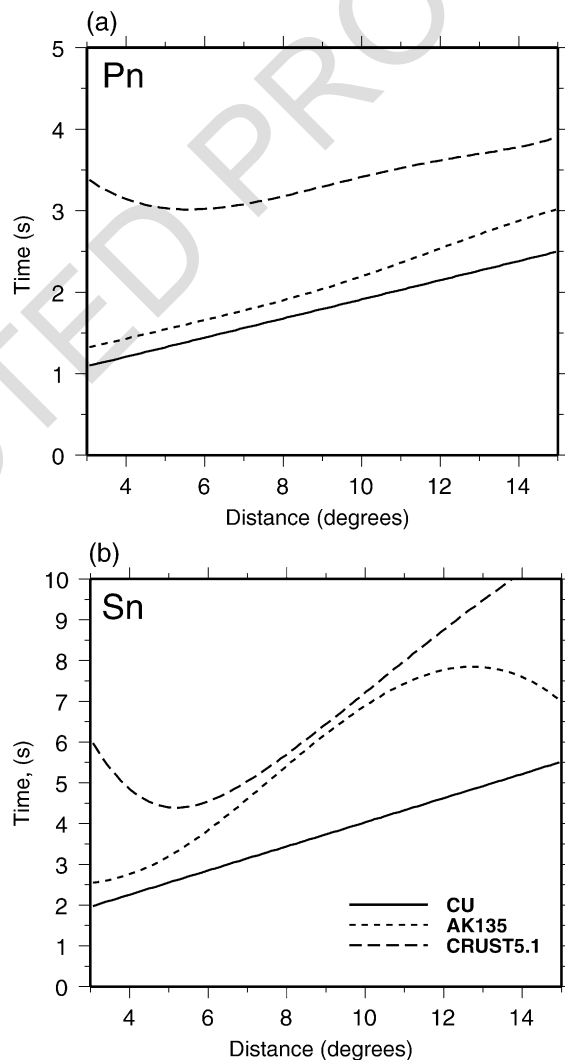


Fig. 10. Smoothed rms misfit versus distance. The solid line is for the CU  $P_n/S_n$  model, the dotted line is for the 1-D model ak135, and the dashed line if for CRUST5.1. (a)  $P_n$ , (b)  $S_n$ .

427 Using Eqs. (1)–(3) and the notation defined in  
 428 Section 3, we define the travel time correction surface  
 429 as follows:

$$\begin{aligned}
 t_{\text{TCS}}(\Delta, \phi) = & t_{\text{m}}(\Delta, \phi) + t_{\text{crust\_sta}}(\Delta, \phi) \\
 & + t_{\text{crust\_evt}}(\Delta, \phi) + \delta t_{\text{sta}} + \delta t_{\text{evt}} \\
 & + \delta t(\Delta) + \delta t_{\text{m}}(\Delta, \phi) - t_{1\text{D}}(\Delta), \quad (8)
 \end{aligned}$$

430 where  $\delta$  and  $\phi$  are distance and azimuth from the  
 432 station to the event, respectively. The prediction from  
 433 a 1-D model,  $t_{1\text{-D}}$ , is subtracted so that the correction  
 434 surface provides a residual relative to this reference.

435 Fig. 8 displays travel time correction surfaces for  
 436 several IMS stations or surrogates. These surfaces are  
 437 for  $P_n$  with surface sources and the station set at the  
 438 local elevation. This differs from correction surfaces  
 439 as they are commonly displayed in which both the  
 440 source and station are on the reference ellipsoid. The  
 441 peak-to-peak anomaly is about 6 s on each and is  
 442 typically twice this value for the  $S_n$  correction sur-  
 443 face. The correction surface for ANTO (Ankara,

Turkey) compares favorably with those for two  
 444 nearby stations in Turkey (KAS, KVT) reported by  
 445 Myers and Schultz (2000) using a different method.  
 446

447 Extending correction surfaces beyond  $15^\circ$  will  
 448 require using a 3-D model to compute the distance  
 449 correction or the use of 3-D tomography.

## 6. Evaluation of the results

450  
 451 Misfits to the entire Eurasian data set for  $P_n$  and  
 452  $S_n$  are shown in Fig. 9. Overall summary statistics  
 453 are presented in Table 1 and rms misfit versus  
 454 distance is summarized in Fig. 10. The standard  
 455 deviation  $\sigma$  reported in Table 1 is computed relative  
 456 to the distance-dependent mean, so that it represents  
 457 the scatter around a trend. In general, short distance  
 458 vertical offsets in Figs. 9 and 10 result in part from  
 459 errors in the crustal model, either in average crustal  
 460 velocities or Moho depths. Errors in the uppermost  
 461 mantle velocities and vertical velocity gradients  
 462 manifest themselves as trends with distance. The

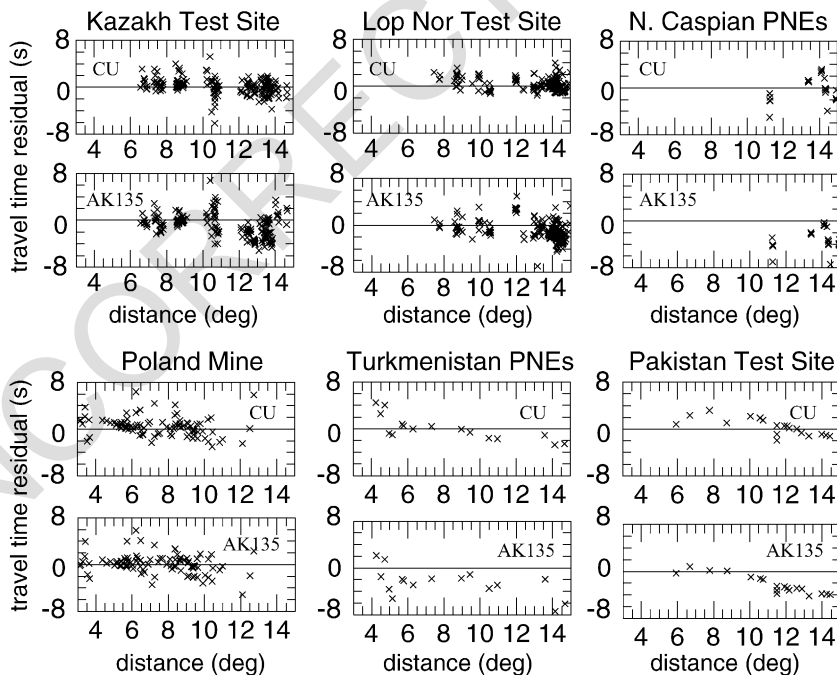


Fig. 11. Misfits to  $P_n$  and  $S_n$  measurements from selected explosions that occurred at six source locations. The upper of each pair of plots for each source location is for the CU  $P_n/S_n$  model and the lower of each pair is for the model ak135. Summary statistics are presented in Table 2.

463 1-D model ak135 does very well for  $P$ . Improvements  
464 afforded by the CU  $P_n/S_n$  model over ak135  
465 are largest at epicentral distances greater than about  
466  $8^\circ$ .  $S$  misfits from ak135, however, exhibit a strong  
467 distance trend, presumably because it is vertically  
468 nearly constant from Moho to about 200 km. Thus,  
469 the misfit trend in  $S$  for ak135 probably results  
470 from an error in the vertical gradient in the upper-  
471 most mantle. For both  $P$  and  $S$ , CRUST5.1 is too  
472 slow in the crust and  $S$  is on average too fast in the  
473 uppermost mantle.

474 The overall rms misfit for the CU  $P_n/S_n$  is 1.6 s  
475 across all of Eurasia for  $P_n$  and approximately twice  
476 this value for  $S_n$ . These misfit statistics appear to be  
477 consistent with those reported by Myers and Schultz  
478 (2000) in a study limited to the neighborhood of the  
479 1991 Racha, Georgia earthquake sequence.

480 The entire Eurasian data set is very noisy and  
481 many locations and origin times are poorly known. A  
482 better estimate of misfit derived from the errors in the  
483 model may come from explosion data in which the  
484 epicenter is well constrained in some cases, although  
485 the origin times may not be. Fig. 11 displays misfits  
486 to data from six explosion regions (three test sites,  
487 two Peaceful Nuclear Explosions (PNEs), and one  
488 large mining district) and Table 2 summarizes these  
489 data. Only explosions with  $m_b \geq 4.6$  as reported in  
490 the PDE catalogue are used. With the exception of a  
491 large mining explosion in southwestern Poland, these  
492 events are not observed at enough stations with  
493 regional phases to have individually constrained  
494 event delays. The overall misfit of the CU  $P_n/S_n$   
495 model to these explosion data is 1.68 s for  $P_n$ , which  
496 is essentially the same as the entire Eurasian data set.  
497 For events at the three test sites, however, misfit is  
498 better than 1.45 s for  $P$  and the fit afforded by ak135  
499 is considerably worse. Again, most of the improve-  
500 ment over ak135 delivered by the CU  $P_n/S_n$  model  
501 comes for paths longer than about  $8^\circ$ . A similar rms  
502 misfit of 1.5 s results from a subset of the complete  
503 Eurasian data set that consists only of events with  
504  $m_b \geq 4.6$  and measurements from events with an  
505 event correction measured at stations with a station  
506 correction. If we remove measurements that fit the  
507 reference model (CRUST5.1) beyond the  $2\sigma$  level,  
508 where  $\sigma$  is the standard deviation of misfit for the  
509 reference model, we find the overall rms misfit for  $P_n$   
510 is about 1.4 s and misfit for  $S_n$  is about 2.8 s.

Table 2

Summary of misfits to explosion data for  $P_n$  displayed in Fig. 11

Location	# Meas.	# Stations	# Events	rms misfit (s)	
				CU $P_n/S_n$	ak135
Kazakh	189	18	42	1.38	2.08
Test Site					
Lop Nor	204	39	19	1.16	1.90
Test Site					
N. Caspian	22	7	6	2.10	3.72
PNEs					
Pakistan	19	18	2	1.43	2.72
Test Site					
Pol Mines	92	92	1	1.72	1.68
Turkmenistan	16	16	1	2.04	3.45
PNEs					
Total	542	–	–	1.68	2.71

511 The rms misfit is probably the best guide to the  
512 accuracy of the correction surfaces. Misfits are a  
513 strong function of epicentral distance, as Fig. 10  
514 shows. Because the distribution of misfit is heavy-  
515 tailed and decidedly non-Gaussian, much of the over-  
516 all misfit comes from bad travel time measurements  
517 that we were unsuccessful in identifying and elimi-  
518 nating prior to inversion. The misfit statistics we  
519 report are, therefore, probably an overestimate of the  
520 error in the predicted travel times.

## 7. Discussion

521  
522 The main purpose of this paper is to assess  $P_n$   
523 and  $S_n$  tomography as a potential means of improv-  
524 ing location capabilities using regional phase data  
525 alone. A full discussion of the velocity anomalies  
526 that appear in the  $P_n$  and  $S_n$  maps, therefore, is  
527 well beyond the intended scope. For greater coher-  
528 ence, however, we mention some of the character-  
529 istics of the estimated maps that agree with shear  
530 velocity anomalies that have emerged from surface  
531 wave dispersion studies (e.g., Shapiro et al., 2000;  
532 Villaseñor et al., 2001). It should be remembered  
533 that  $P_n$  and  $S_n$  maps are of velocities right at the  
534 top of the mantle and are mute about vertical  
535 velocity variations that are revealed by 3-D models.  
536 We will limit this discussion to Central Asia.

537 The old, stable cratons north of the Alpine–  
538 Himalayan orogenic belt are characterized by high



539 upper-mantle  $P_n$  and  $S_n$  velocities. High velocities are  
540 also found beneath the Indian shield, the southern  
541 Tibetan Plateau, and the Tarim Basin. While high  
542 velocities in the upper mantle are usually interpreted  
543 as an indication of old, cold, thick lithospheric  
544 blocks, the structures associated with low velocity  
545 anomalies are more difficult to interpret. Large low  
546 velocity anomalies are associated with young, exten-  
547 sional plate boundaries, such as the Red Sea. The low  
548 velocity anomaly beneath central and northern Tibet  
549 has received a great deal of attention (see Molnar,  
550 1988) because of its implications for the origin and  
551 mechanism for the formation of the Tibetan plateau.  
552 However, although present in the  $P_n$  and  $S_n$  maps, it  
553 is not one of the most prominent negative anomalies  
554 in magnitude or in extent. Based upon the presence  
555 of this low velocity region and other evidence (e.g.,  
556 widespread Quaternary volcanism and inefficient  $S_n$   
557 propagation), Molnar et al. (1993) proposed that the  
558 high-velocity Indian lithosphere has not been under-  
559 thrust beneath the Tibetan Plateau, and that crustal  
560 thickening has occurred by north–south shortening of  
561 the southern Eurasian crust.

562 One of the most prominent upper-mantle low  
563 velocity regions is located in the Middle East,  
564 extending from Turkey to Iran and western Afgha-  
565 nistan. This low velocity anomaly coincides with  
566 the Turkish–Iranian continental plateau, formed by  
567 the collision between the Arabian and Eurasian  
568 plates. This collision is the result of the closing  
569 of the Neo-Tethys Ocean by northward subduction  
570 of oceanic lithosphere beneath Eurasia. In Iran and  
571 western Afghanistan, the low velocity anomaly is  
572 bounded to the south by high velocities, part of the  
573 Arabian plate. This low velocity anomaly is prom-  
574 inent in other  $P_n$  tomography studies (e.g., Hearn  
575 and Ni, 1994), and is also coincident with a region  
576 of high S-wave attenuation (Kadinsky-Cade et al.,  
577 1981) and Neogene volcanism (Kazmin et al.,  
578 1986). The combination of these observations sug-  
579 gests a hot or perhaps partially molten uppermost  
580 mantle beneath the Turkish–Iranian Plateau. This  
581 anomalously hot upper mantle could be a remnant  
582 of the back-arc extensional regime that dominated  
583 this region from the Jurassic to the Neogene  
584 (Dercourt et al., 1986). The presence of hot, molten  
585 upper mantle weakens the lithosphere, allowing  
586 larger deformation associated with the Arabian

587 plate–Eurasia collision. This results in the observed  
588 diffuse intraplate seismicity that extends well to the  
589 north of the plate boundary delineated by the  
590 Zagros Main Thrust. Furthermore, the buoyancy  
591 associated with hot upper mantle, combined with  
592 the buoyancy due to the deep continental roots in  
593 the region, can contribute to maintain the high  
594 topography of the plateau.

595 Another significant upper-mantle low velocity  
596 anomaly is centered in western Mongolia, WSW of  
597 lake Baikal. The central part of this anomaly coincides  
598 with the Hangay Dome area of central Mongolia. The  
599 Hangay Dome is characterized by recent uplift, dif-  
600 fuse extension and regionally upwarped topography  
601 (Cunningham, 1988). This is also a region of recent  
602 Cenozoic volcanism and high heat flow (with a  
603 maximum of approximately 80 mW/m<sup>2</sup>). There is  
604 remarkable agreement between the shape of the  
605 velocity anomaly and the heat flow anomaly (Fig.  
606 5a of Cunningham, 1988). This region in Mongolia  
607 has been interpreted to overlie a mantle plume or  
608 asthenospheric diapir, which is associated with rifting  
609 in Lake Baikal (Windley and Allen, 1993). Lake  
610 Baikal is located at the boundary between the Mon-  
611 golian plateau and the Siberian Craton, which is  
612 consistent with the marked contrast between low  
613 and high velocities observed in our  $P_n$  and  $S_n$  maps.  
614 Irrespective of its cause, deformation due to the  
615 presence of the mantle plume or asthenospheric diapir  
616 manifests in high seismic activity in Western Mongo-  
617 lia, which has been the site of some of the largest  
618 intraplate earthquakes recorded during this century  
619 (i.e., 1905 and 1957).

## 8. Conclusions

620  
621 The method for producing  $P_n$  and  $S_n$  maps with  
622 associated parametric corrections effectively summa-  
623 rizes the information in the large groomed ISC/NEIC  
624 data base for epicentral distances from about 3° to  
625 15°. The  $P_n$  and  $S_n$  maps correlate well with other  
626 high resolution information about structural variations  
627 in the uppermost mantle. In particular, these maps  
628 produce relatively high resolution images of low  
629 velocity anomalies in tectonically deformed regions  
630 across the continent. These include anomalies across  
631 central and southern Asia and the Middle East that

632 extend along a tortuous path from Turkey in the west  
633 to Lake Baikal in the east. These anomalies are related  
634 to the closing of the Neo-Tethys Ocean and the  
635 collision of India with Asia. The uppermost mantle  
636 beneath the Pacific Rim back-arc is also very slow,  
637 presumably due to upwelling that results from back-  
638 arc spreading, as is the Red Sea rift, the Tyrrhenian  
639 Sea and other regions undergoing active extension.

640 The travel time correction surfaces computed from  
641 the CU  $P_n/S_n$  model appear to be robust and fit the  
642 data with low levels of bias at epicentral distances  
643 from  $3^\circ$  to  $15^\circ$ . Overall rms misfits across Eurasia for  
644  $P_n$  are  $\sim 1.6$  s and for  $S_n$   $\sim 3.2$  s, are better for data  
645 subsets chosen for their quality (e.g., explosions, large  
646 magnitude events, independent information about epi-  
647 center location and/or origin time), and exhibit a  
648 strong, nearly linear distance trend. These misfits are  
649 considerably better than those produced by ak135 and  
650 CRUST5.1, although ak135 fits the  $P$  data remarkably  
651 well for a 1-D model. The correction surfaces pre-  
652 sented here provide a reference for 3-D models to  
653 match and extend.

654 Although the method described here appears to  
655 produce reliable travel time correction surfaces, there  
656 are greater problems in estimating  $P_n$  and  $S_n$  reliably  
657 due to trade-offs between the estimated tomographic  
658 map and the parametric corrections. Some of these  
659 trade-offs can be ameliorated in the future if a 3-D  
660 model is used as the reference model, which will  
661 allow the horizontal ray approximation to be broken.  
662 Indeed, it is likely that our  $P_n$  model fits the data only  
663 marginally better than the 1-D model ak135 because  
664 a single distance correction is inadequate to model  
665 ray penetration into the upper mantle, which can be  
666 highly variable, as Fig. 3b indicates. Getting the  
667 vertical velocity derivative right may be more impor-  
668 tant in predicting regional travel times than mapping  
669 lateral variations. Recent models, such as those of  
670 Ekström and Dziewonski (1997), Villaseñor et al.  
671 (2001), and Shapiro et al. (2000), are providing new  
672 information about the vertical velocity gradient in the  
673 uppermost mantle which controls the depth of pene-  
674 tration and, hence, a large fraction of the travel time  
675 of regionally propagating phases. In addition, to  
676 extend travel time correction surfaces beyond  $15^\circ$   
677 will require a 3-D model to predict the ray paths.

678 The final proof of the effectiveness of the method  
679 described here will be the relocation of ground-truth

events. The agreement between observed and pre-  
dicted travel times for Eurasian explosions is encour-  
aging, but rigorous tests to determine the extent to  
which the correction surfaces will improve regional  
location capabilities define a crucial remaining hurdle.

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